

**Reunion IMNC/SHFJ  
24 Fevrier 2012**

# Dielectric-dielectric surface

```
<surface name="MySurface" type="dielectric_dielectric" sigmaalpha="20.0" finish="ground">
  <propertystable>
    <propertyvector name="SPECULARLOBECONSTANT" energyunit="eV">
      <ve energy="2.34" value="0.0"/>
      <ve energy="1.97" value="0.0"/>
    </propertyvector>
    <propertyvector name="SPECULARSPIKECONSTANT" energyunit="eV">
      <ve energy="2.34" value="0.0"/>
      <ve energy="1.97" value="0.0"/>
    </propertyvector>
    <propertyvector name="BACKSCATTERCONSTANT" energyunit="eV">
      <ve energy="2.34" value="0.0"/>
      <ve energy="1.97" value="0.0"/>
    </propertyvector>
  <!--
    <propertyvector name="REFLECTIVITY" energyunit="eV">
      <ve energy="2.34" value="1.0"/>
      <ve energy="1.97" value="1.0"/>
    </propertyvector-->
  </propertystable>
</surface>
```

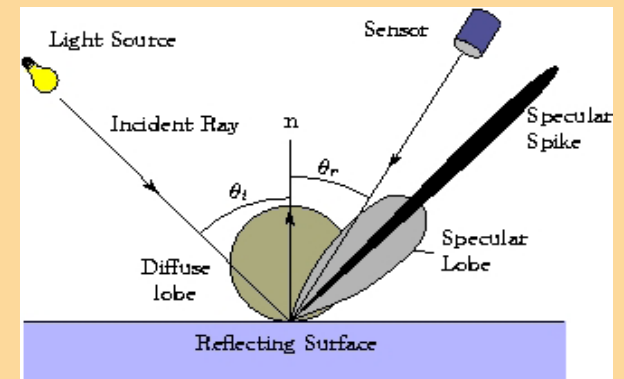
Reflection about the normal of a microfacet

Reflection about the AVERAGE surface normal

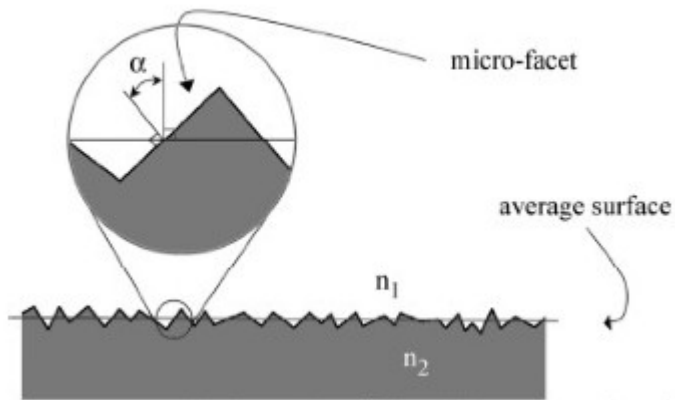
Several reflections with the ultimate result of exact back-scatter

Example of 100% Lambertian Reflection

- 4 parameters: Reflectivity, Specular Lobe, Specular Spike, Back-scatter (Lambertian is implicit)
- Default Reflectivity is 1.
- Surface finish: polished and ground
- Polished interface: no surface needs to be specified (only 2 media Rindex)
- Ground interface: If not 100% Lambertian (diffuse), you can simulate specular and ultimate back-scatter.
- Not all rough surfaces are perfect lambertian reflectors, but it is a good approximation if surface characteristics are unknown (Geant4 Developer guide).



# SigmaAlpha dependence



The surface normals micro-facets are distributed around the average surface normal, following a normal distribution with standard deviation:

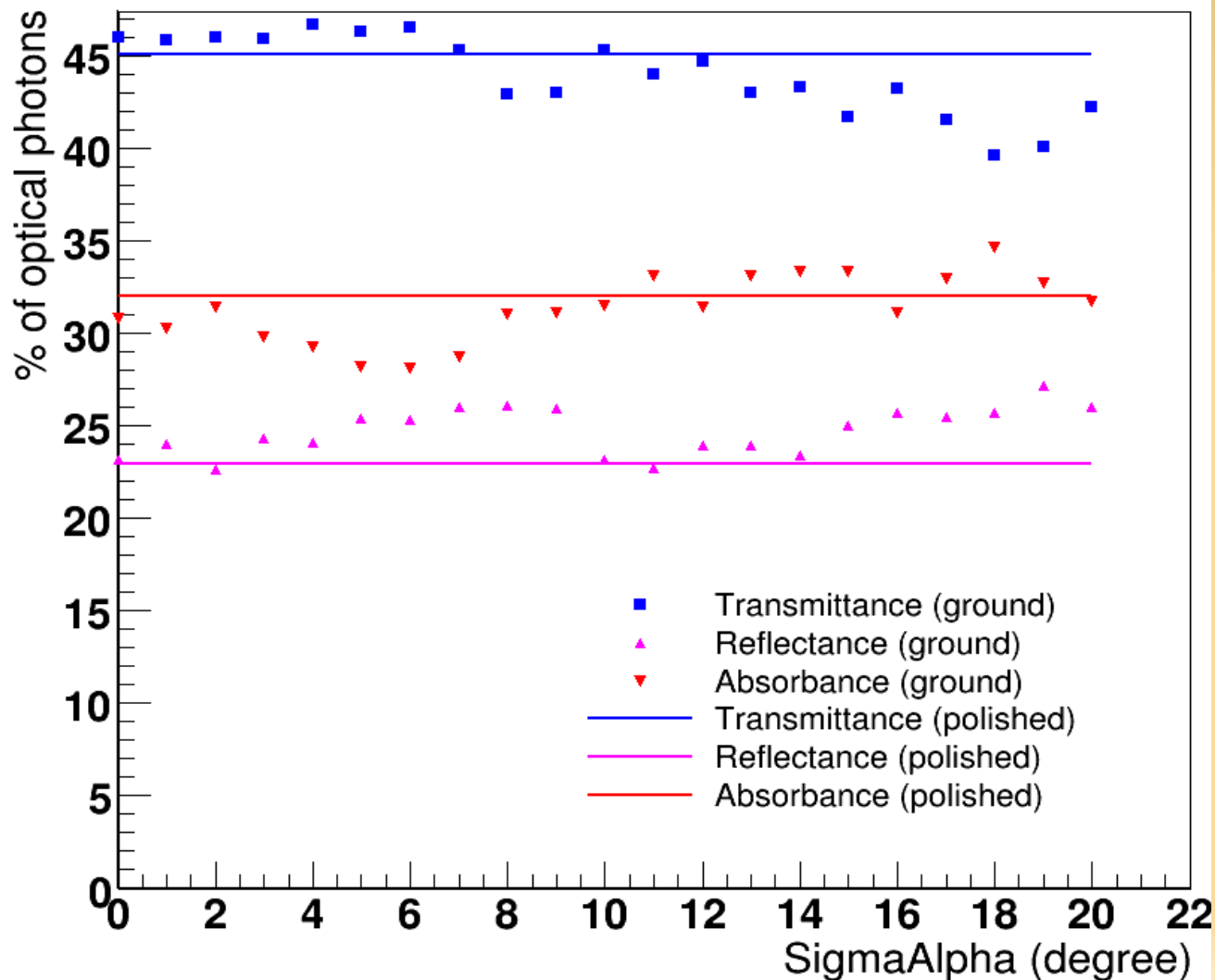
$$\sigma_{\alpha}$$

Difference is ~ 5%

5.5% T

4.3% R

3.9% A



# hgate Physics Models

- Version 3 du code hGate (J. Bert)
- Physics Models for EM: Standard and Livermore (low energy electron and gamma)
  - Livermore models for  $\gamma$ ,  $e^-$  below 1 GeV
  - Standard models above 1 GeV
- Standard Compton scattering: interaction of a high energy photon with an electron ==> Klein-Nishina formula: differential cross section.
- Standard Photo Electric Effect (ejection of an electron from a material after a photon has been absorbed by that material): Parameterized formula (Biggs et al – Sandia Tables)

# Livermore Low Energy Model

- Based on the Livermore database (experimental data parameterisations) developed by the Lawrence Livermore National Laboratory.
- Rayleigh Scattering: Angular distribution  $\Phi(E, \theta) = [1 + \cos^2(\theta)] \cdot F^2(q)$ , where  $F(q)$  is the energy-dependent form factor (LLNL database).
- Compton Scattering: Energy distribution of the scattered photon according to Klein-Nishina formula multiplied by scattering functions (LLNL database).
- PhotoElectric Effect: Subshell from which the electron is emitted is selected according to the cross sections of the sub-shells, transition probabilities from LLNL database.

# v03\_hgate\_livermore\_standard: Main Code

## ■ mc\_main.cu

- Number of required simulated particles
- Physics list choice = Rayleigh, Compton, PhotoElectric
- Water box is built
- Physics model choice = Standard or Livermore
  - NB: there is no Standard Model Rayleigh.
- Woodcock tracking (also called delta tracking):
  - Material cross-sections vary within a geometric region
  - One common approach: Breaking the flight path into short segments within which the cross-section is assumed constant. ==> expensive calculations.
  - Delta-tracking approach: Uses a fictitious cross-section  $\Sigma_{\max}$  (the largest value expected during the particle flight). A trial flight distance is sampled using that cross section and the particle is moved that distance.
  - $P = \Sigma(s) / \Sigma_{\max}$  is determined. With probability  $P$ , the flight distance is accepted; with probability  $1-P$ , it is rejected, and the entire procedure is repeated.

# v03\_hgate\_livermore\_standard: Functions

## ■ mc\_fun.cu

- Functions from CLHEP: loglog interpolation, linear interpolation, deflect particle (due to Compton or Rayleigh scatter)
- Random Number generator BRENT xor256
- Livermore: Rayleigh, Compton and PhotoElectric cross sections per Atom
- Standard: Compton and PhotoElectric cross sections per Atom
- Scatter Factors (Livermore): Rayleigh and Compton
- Functions: Compton scatter (STD and Livermore), Rayleigh Scatter (Livermore only), Total Compton and Photoelectric cross sections for a given material (STD and Livermore), Rayleigh Total cross section for a given material (Livermore only)
- Compton and Rayleigh Element Selector: Z material mixture
- Sources: Fictitious tracking (STD and Livermore), Interaction (physics and new particle direction), Phase space (which particle will be reloaded in stack), Open cross sections Tables.

# v03\_hgate\_livermore\_standard: Materials, Cross-Sections, Data

## List of index materials

- |               |               |
|---------------|---------------|
| 0 – Air       | 8 – Intestine |
| 1 – Water     | 9 – Spleen    |
| 2 – Body      | 10 – Blood    |
| 3 – Lung      | 11 – Liver    |
| 4 – Breast    | 12 – Kidney   |
| 5 – Heart     | 13 – Brain    |
| 6 – SpineBone | 14 - Pancreas |
| 7 - RibBone   |               |

## Data files (binary)

Compton (x-sec and SF) Livermore

PhotoElectric (x-sec)

Rayleigh (x-sec and SF) Livermore

mc\_cst.cu

Constants, Materials

Some cross-sections

Scatter Factors (Compton, Rayleigh)

PhotoElectric:

Atomic number (Z)/ Atomic Mass (A)

Ionization Potentials

Sandia Tables (Z=1 to 30)

Atomic Number Density for each  
element of materials

Atom Bind Energy

Doppler Profiles for Compton ...



# Notre status @ CCRT

- Le groupe DSV n'a pas d'accès au GPU sur la machine Titane (48 serveurs GPU Tesla Nvidia destinés aux utilisateurs GENCI)
- Contact: Christine Menache
  - Titane-GPU: demande Dari pour le 2ème semestre
  - Curie-GPU: demande Preparatory Access de type B qui dure 6 mois et donne accès à ~60 000h GPU (deadline 01.03.12).
- BACKUP plan: Utilisation de la machine Cesium (chaque nœud est équipé d'une carte GPU Nvidia) [tests uniquement]

# Environnement et Compilation

- Environnement de travail CUDA sur Cesium

```
#!/bin/bash
```

```
#Chargement des modules GPU  
module load cuda/4.0
```

```
#Variable MPI_HOME  
export MPI_HOME=`which mpicc | sed 's/\sbin/mpicc//g'`
```

- Compilation code

```
nvcc -o mc_main mc_main.cu -arch compute_20 -use_fast_math
```

Aucun Probleme!

# Execution du code sur Cesium

- Les calculs sur les cartes graphiques se font en batch uniquement

```
#!/bin/bash
#MSUB -r slinky
#MSUB -N 1 # nombre de noeuds a allouer
#MSUB -n 1 # nombre de processeurs a allouer
##MSUB -T 1800
#MSUB -q test
#MSUB -o slinky.o
#MSUB -e slinky.e
set -x
cd $BRIDGE_MSUB_PWD
#mpirun -np 1 /applications/cuda-4.0/sdk/C/bin/linux/release/deviceQuery
mpirun -np 1 ./mc_main 10
```

argument: Random Seed

340459.cesium21	15	pool11	prod@cesium21	2575	ROO	200	86400	200	86400	07	7168	1	1
340460.cesium22	slinky	vcuplov	test@cesium22	0	PEN	0	1800	0	1800	0	7168	0	1
340461.cesium23	ctrl	r241	prod@cesium23	0	PEN	0	86400	0	86400	0	7168	0	1

```
Tot particles simulated: 1000 - 1000
Running time is 0.042941 s - tot export 0
```

- Differente architecture: Fermi (2.0) (julien) versus Quadro (1.3)
- memoire insuffisante? Non :(
- en contact (permanent) avec le prof de la formation GPU au CCC :)

# Resultats

- Cesium ne fonctionne pas bien en batch
- Le code tourne (6s) sur Titane
- Pour continuer à travailler sur Cesium, il faut utiliser visuportal (bureau virtuel) et travailler en interactif.

